Entropy and quantum field theory

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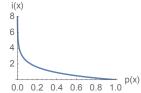


Entropy and information

[Claude Shannon (1948)]

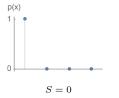
- ullet consider a random variable x with probability distribution p(x)
- ullet information content or "surprise" associated with outcome x

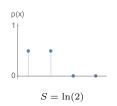
$$i(x) = -\ln p(x)$$

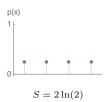


• Entropy is expectation value of information content

$$S = \langle i(x) \rangle = -\sum p(x) \ln p(x)$$







Entropy at thermal equilibrium

- ullet micro canonical ensemble: maximal entropy S for given conserved quantities E,N in given volume V
- universality at equilibrium
- starting point for development of thermodynamics ...

$$S(E, N, V),$$

$$dS = \frac{1}{T}dE - \frac{\mu}{T}dN + \frac{p}{T}dV$$

• ... grand canonical ensemble with density operator ...

$$\rho = \frac{1}{Z}e^{-\frac{1}{T}(H-\mu N)}$$

• ... Matsubara formalism for quantum fields ...

Ideal fluid dynamics

• thermal equilibrium

$$T^{\mu\nu} = \epsilon u^{\mu}u^{\nu} + p(u^{\mu}u^{\nu} + g^{\mu\nu}), \qquad N^{\mu} = nu^{\mu}, \qquad s^{\mu} = su^{\mu}$$

- ullet fluid velocity u^μ
- ullet thermodynamic equation of state $p(T,\mu)$ with $dp=sdT+nd\mu$
- local thermal equilibrium approximation: $u^{\mu}(x)$, T(x), $\mu(x)$
- neglect gradients: lowest order of a derivative expansion
- ullet evolution of $u^{\mu}(x)$, T(x) and $\mu(x)$ from conservation laws

$$\nabla_{\mu} T^{\mu\nu}(x) = 0, \qquad \nabla_{\mu} N^{\mu}(x) = 0.$$

entropy current also conserved

$$\nabla_{\mu} s^{\mu}(x) = 0.$$

$Out ext{-}of ext{-}equilibrium$

- quantum field theory out-of-equilibrium is less well understood
- interesting topic of current research
- is non-equilibrium dynamics also governed by information?
- approach to equilibrium
- universality

Entropy in quantum theory

[John von Neumann (1932)]

$$S = -\mathsf{Tr}\rho\ln\rho$$

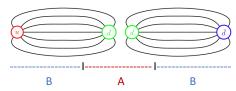
- ullet based on the quantum density operator ho
- for pure states $\rho = |\psi\rangle\langle\psi|$ one has S=0
- for mixed states $\rho = \sum_{j} p_{j} |j\rangle\langle j|$ one has $S = -\sum_{j} p_{j} \ln p_{j} > 0$
- unitary time evolution conserves entropy

$$-{\rm Tr}(U\rho U^\dagger)\ln(U\rho U^\dagger) = -{\rm Tr}\rho\ln\rho \qquad \to \qquad S = {\rm const.}$$

• global characterization of quantum state

Entropy and entanglement

ullet consider a split of a quantum system into two A+B



ullet reduced density operator for system A

$$\rho_A = \mathsf{Tr}_B\{\rho\}$$

entropy associated with subsystem A

$$S_A = -\mathsf{Tr}_A \{ \rho_A \ln \rho_A \}$$

- pure **product** state $\rho = \rho_A \otimes \rho_B$ leads to $S_A = 0$
- pure entangled state $\rho \neq \rho_A \otimes \rho_B$ leads to $S_A > 0$
- S_A is called **entanglement entropy**

Classical statistics

- ullet consider system of two random variables x and y
- ullet joint probability p(x,y) , joint entropy

$$S = -\sum_{x,y} p(x,y) \ln p(x,y)$$

- reduced or marginal probability $p(x) = \sum_{y} p(x, y)$
- reduced or marginal entropy

$$S_x = -\sum_x p(x) \ln p(x)$$

 one can prove: joint entropy is greater than or equal to reduced entropy

$$S \ge S_x$$

• globally pure state S=0 is also locally pure $S_x=0$

Quantum statistics

- ullet consider system with two subsystems A and B
- ullet combined state ho , combined or full entropy

$$S = -\mathsf{Tr}\{\rho \ln \rho\}$$

- reduced density matrix $\rho_A = \text{Tr}_B\{\rho\}$
- reduced or entanglement entropy

$$S_A = -\mathsf{Tr}_A \{ \rho_A \ln \rho_A \}$$

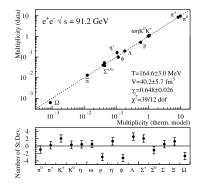
• for quantum systems entanglement makes a difference

$$S \ngeq S_A$$

- coherent information $I_{B \setminus A} = S_A S$ can be positive!
- globally pure state S=0 can be locally mixed $S_A>0$

The thermal model puzzle

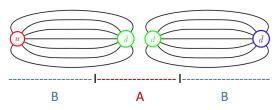
- \bullet elementary particle collision experiments such as $e^+\ e^-$ collisions show thermal-like features
- particle multiplicities well described by thermal model



[Becattini, Casterina, Milov & Satz, EPJC 66, 377 (2010)]

- conventional thermalization by collisions unlikely
- alternative explanations needed

$QCD\ strings$



- particle production from QCD strings
- e. g. Lund model (Pythia)
- different regions in a string are entangled
- ullet subinterval A is described by reduced density matrix of mixed form

$$\rho_A = \mathsf{Tr}_B \rho$$

characterization by entanglement entropy

$$S_A = -\text{Tr}\left\{\rho_A \ln(\rho_A)\right\}$$

• could this lead to thermal-like effects?

$Microscopic\ model$

QCD in 1+1 dimensions described by 't Hooft model

$$\mathcal{L} = -\bar{\psi}_i \gamma^{\mu} (\partial_{\mu} - ig\mathbf{A}_{\mu}) \psi_i - m_i \bar{\psi}_i \psi_i - \frac{1}{2} \mathrm{tr} \, \mathbf{F}_{\mu\nu} \mathbf{F}^{\mu\nu}$$

- ullet fermionic fields ψ_i with sums over flavor species $i=1,\dots,N_f$
- ullet SU (N_c) gauge fields ${f A}_{\mu}$ with field strength tensor ${f F}_{\mu
 u}$
- gluons are not dynamical in two dimensions
- ullet gauge coupling g has dimension of mass
- non-trivial, interacting theory, cannot be solved exactly
- \bullet spectrum of excitations known for $N_c \to \infty$ with $g^2 N_c$ fixed ['t Hooft (1974)]

$Schwinger\ model$

• QED in 1+1 dimension

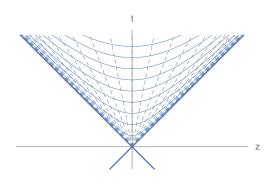
$$\mathscr{L} = -\bar{\psi}_i \gamma^{\mu} (\partial_{\mu} - iqA_{\mu}) \psi_i - m_i \bar{\psi}_i \psi_i - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

- geometric confinement
- U(1) charge related to string tension $q = \sqrt{2\sigma}$
- for single fermion one can bosonize theory exactly [Coleman, Jackiw, Susskind (1975)]

$$S = \int d^2x \sqrt{g} \left\{ -\frac{1}{2} g^{\mu\nu} \partial_{\mu} \phi \partial_{\nu} \phi - \frac{1}{2} M^2 \phi^2 - \frac{m q e^{\gamma}}{2\pi^{3/2}} \cos \left(2\sqrt{\pi}\phi + \theta\right) \right\}$$

- ullet Schwinger bosons are dipoles $\phi \sim ar{\psi} \psi$
- mass is related to U(1) charge by $M=q/\sqrt{\pi}=\sqrt{2\sigma/\pi}$
- ullet massless Schwinger model m=0 leads to free bosonic theory

$Expanding\ string\ solution$



- ullet external quark-anti-quark pair on trajectories $z=\pm t$
- \bullet coordinates: Bjorken time $\tau = \sqrt{t^2 z^2},$ rapidity $\eta = \operatorname{arctanh}(z/t)$
- $\bullet \ \mathrm{metric} \ ds^2 = -d\tau^2 + \tau^2 d\eta^2$
- \bullet symmetry with respect to longitudinal boosts $\eta \to \eta + \Delta \eta$

Coherent field evolution

ullet Schwinger boson field depends only on au

$$\bar{\phi} = \bar{\phi}(\tau)$$

equation of motion

$$\partial_{\tau}^{2}\bar{\phi} + \frac{1}{\tau}\partial_{\tau}\bar{\phi} + M^{2}\bar{\phi} = 0.$$

• Gauss law: electric field $E=q\phi/\sqrt{\pi}$ must approach the U(1) charge of the external quarks $E\to q_{\rm e}$ for $\tau\to 0_+$

$$\bar{\phi}(\tau) \to \frac{\sqrt{\pi}q_{\mathsf{e}}}{q} \qquad (\tau \to 0_+)$$

solution of equation of motion [Loshaj, Kharzeev (2011)]

$$ar{\phi}(au) = rac{\sqrt{\pi}q_{\mathsf{e}}}{q}J_0(M au)$$

Gaussian states

- theories with quadratic action typically have Gaussian density matrix
- fully characterized by field expectation values

$$\bar{\phi}(x) = \langle \phi(x) \rangle, \qquad \bar{\pi}(x) = \langle \pi(x) \rangle$$

and connected two-point correlation functions, e. g.

$$\langle \phi(x)\phi(y)\rangle_c = \langle \phi(x)\phi(y)\rangle - \bar{\phi}(x)\bar{\phi}(y)$$

ullet if ho is Gaussian, also reduced density matrix ho_A is Gaussian

Entanglement entropy for Gaussian state

ullet entanglement entropy of Gaussian state in region A [Berges, Floerchinger, Venugopalan, 1712.09362]

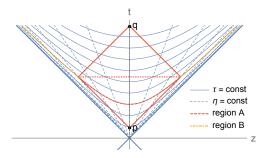
$$S_A = \frac{1}{2} \operatorname{Tr}_A \left\{ D \ln(D^2) \right\},\,$$

- ullet operator trace over region A only
- matrix of correlation functions

$$D(x,y) = \begin{pmatrix} -i\langle\phi(x)\pi(y)\rangle_c & i\langle\phi(x)\phi(y)\rangle_c \\ -i\langle\pi(x)\pi(y)\rangle_c & i\langle\pi(x)\phi(y)\rangle_c \end{pmatrix}.$$

- involves connected correlation functions of field $\phi(x)$ and canonically conjugate momentum field $\pi(x)$
- ullet expectation value $ar{\phi}$ does not appear explicitly
- ullet coherent states and vacuum have equal entanglement entropy S_A

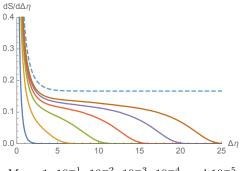
$Rapidity\ interval$



- ullet consider rapidity interval $(-\Delta\eta/2,\Delta\eta/2)$ at fixed Bjorken time au
- entanglement entropy does not change by unitary time evolution with endpoints kept fixed
- can be evaluated equivalently in interval $\Delta z = 2\tau \sinh(\Delta\eta/2)$ at fixed time $t = \tau \cosh(\Delta\eta/2)$
- need to solve eigenvalue problem with correct boundary conditions

Bosonized massless Schwinger model

- entanglement entropy understood numerically for free massive scalars [Casini, Huerta (2009)]
- entanglement entropy density $dS/d\Delta\eta$ for bosonized massless Schwinger model $(M=\frac{q}{\sqrt{\pi}})$



 $M\tau=1,\,10^{-1},\,10^{-2},\,10^{-3},\,10^{-4},\,{\rm and}\,\,10^{-5}$

[Berges, Floerchinger, Venugopalan (2017)]

Conformal limit

• for M au o 0 one has conformal field theory limit [Holzhey, Larsen, Wilczek (1994); Calabrese, Cardy (2004)]

$$S(\Delta z) = \frac{c}{3} \ln \left(\Delta z / \epsilon \right) + \text{constant}$$

with small length ϵ acting as UV cutoff

here this implies

$$S(\tau,\Delta\eta) = \frac{c}{3} \ln \left(2\tau \sinh(\Delta\eta/2)/\epsilon \right) + {\rm constant}$$

- ullet conformal charge c=1 for free massless scalars or Dirac fermions
- additive constant not universal but entropy density is

$$\begin{split} \frac{\partial}{\partial \Delta \eta} S(\tau, \Delta \eta) = & \frac{c}{6} \mathrm{coth}(\Delta \eta / 2) \\ \rightarrow & \frac{c}{6} \qquad (\Delta \eta \gg 1) \end{split}$$

ullet entropy becomes extensive in $\Delta\eta$!

Universal entanglement entropy density

 for very early times "Hubble" expansion rate dominates over masses and interactions

$$H = \frac{1}{\tau} \gg M = \frac{q}{\sqrt{\pi}}, m$$

- theory dominated by free, massless fermions
- universal entanglement entropy density

$$\frac{dS}{d\Delta\eta} = \frac{c}{6}$$

with conformal charge \boldsymbol{c}

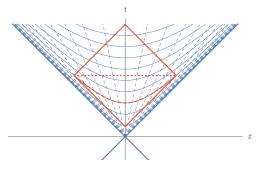
• for QCD in 1+1 dimensions (gluons not dynamical)

$$c = N_c \times N_f$$

• from fluctuating transverse coordinates (Nambu-Goto action)

$$c = N_c \times N_f + 2 \approx 9 + 2 = 11$$

Modular or entanglement Hamiltonian



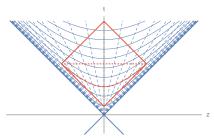
• conformal field theory [Casini, Huerta, Myers (2011), Arias, Blanco, Casini, Huerta (2017), see also Candelas, Dowker (1979)]

$$\rho_A = \frac{1}{Z_A} e^{-K}, \qquad Z_A = \operatorname{Tr} e^{-K}$$

• modular or entanglement Hamiltonian local expression

$$K = \int_{\Sigma} d\Sigma_{\mu} \, \xi_{\nu}(x) \, T^{\mu\nu}(x)$$

Time-dependent temperature



- ullet energy-momentum of excitations around coherent field $T^{\mu
 u}(x)$
- combination of fluid velocity and temperature $\xi^{\mu}(x) = \frac{u^{\mu}(x)}{T(x)}$
- fluid velocity in τ -direction & time-dependent temperature [Berges, Floerchinger, Venugopalan (2017)]

$$T(\tau) = \frac{\hbar}{2\pi\tau}$$

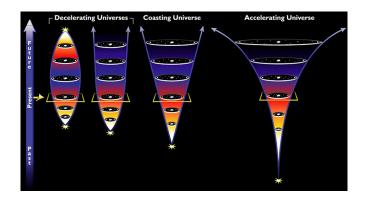
- Entanglement between different rapidity intervals alone leads to local thermal density matrix at very early times!
- Hawking-Unruh temperature in Rindler wedge $T(x) = \frac{\hbar c}{2\pi x}$

Physics picture

- alternative derivation via mode functions & Bogoliubov transforms [Berges, Floerchinger, Venugopalan, 1712.09362]
- coherent state vacuum at early time contains entangled pairs of quasi-particles with opposite wave numbers
- on finite rapidity interval $(-\Delta\eta/2,\Delta\eta/2)$ in- and out-flux of quasi-particles with thermal distribution via boundaries
- ullet technically **limits** $\Delta \eta \to \infty$ and $M au \to 0$ do not commute
 - $\Delta\eta o \infty$ for any finite M au gives pure state
 - M au o 0 for any finite $\Delta \eta$ gives thermal state with $T=1/(2\pi au)$

Entanglement dynamics in cold atom experiments

- entanglement can be directly accessed in cold atom experiments [Oberthaler group, Greiner group]
- expanding geometries can be realized by interplay of
 - longitudinal expansion
 - time dependent change of sound velocity $v_s(t)$
 - time dependent gap or mass $M^2(t)$



Dissipation

• dissipation can be defined as (effective) entropy generation

$$\frac{d}{dt}S > 0$$

• for extensive entropy $S=\int_{\Sigma}d\Sigma_{\mu}s^{\mu}$ one has locally

$$\nabla_{\mu}s^{\mu} > 0$$

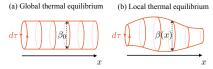
- related to effective loss of information
- second law of thermodynamics: entropy gets produced, not destroyed
- local dissipation entanglement generation (?)

Dissipation and the quantum effective action

- dissipation usually discussed on the level of equations of motion
- one would like to have a formulation in terms of an effective action
 - fluctuations & correlation functions
 - renormalization
 - effective field theories
 - coupling to gravity
- one possibility: Schwinger-Keldysh double time path formalism
- another possibility: analytic continuation of the 1PI effective action [Floerchinger, JHEP 1609, 099 (2016)]

Local equilibrium & partition function

[Floerchinger, JHEP 1609, 099 (2016)]



• local equilibrium with T(x) and $u^{\mu}(x)$

$$\beta^{\mu}(x) = \frac{u^{\mu}(x)}{T(x)}$$

similarity between local density matrix and translation operator

$$e^{\beta^{\mu}(x)\mathscr{P}_{\mu}} \longleftrightarrow e^{i\Delta x^{\mu}\mathscr{P}_{\mu}}$$

represent partition function as functional integral with periodicity

$$\phi(x^{\mu} - i\beta^{\mu}(x)) = \pm \phi(x^{\mu})$$

ullet partition function Z[J], Schwinger functional W[J] in Euclidean

$$Z[J] = e^{W_E[J]} = \int D\phi \, e^{-S_E[\phi] + \int_x J\phi}$$

One-particle irreducible or quantum effective action

ullet in Euclidean domain $\Gamma[\phi]$ defined by Legendre transform

$$\Gamma_E[\Phi] = \int_x J_a(x)\Phi_a(x) - W_E[J]$$

with expectation values

$$\Phi_a(x) = \frac{1}{\sqrt{g}(x)} \frac{\delta}{\delta J_a(x)} W_E[J]$$

Euclidean field equation

$$\frac{\delta}{\delta \Phi_a(x)} \Gamma_E[\Phi] = \sqrt{g}(x) J_a(x)$$

resembles classical equation of motion for J=0

• need analytic continuation to obtain a viable equation of motion

Two-point functions

homogeneous background field and global equilibrium

$$\beta^{\mu} = \left(\frac{1}{T}, 0, 0, 0\right)$$

propagator and inverse propagator

$$\frac{\delta^2}{\delta J_a(-p)\delta J_b(q)} W_E[J] = G_{ab}(p) (2\pi)^4 \delta^{(4)}(p-q)$$
$$\frac{\delta^2}{\delta \Phi_a(-p)\delta \Phi_b(q)} \Gamma_E[\Phi] = P_{ab}(p) (2\pi)^4 \delta^{(4)}(p-q)$$

from definition of effective action

$$\sum_{b} G_{ab}(p) P_{bc}(p) = \delta_{ac}$$

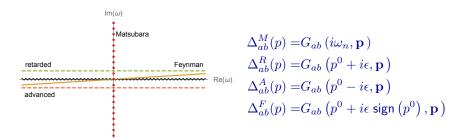
Spectral representation

Källen-Lehmann spectral representation

$$G_{ab}(\omega, \mathbf{p}) = \int_{-\infty}^{\infty} dz \, \frac{\rho_{ab}(z^2 - \mathbf{p}^2, z)}{z - \omega}$$

with $\rho_{ab} \in \mathbb{R}$

- \bullet correlation functions can be analytically continued in $\omega = -u^\mu p_\mu$
- ullet branch cut or poles on real frequency axis $\omega \in \mathbb{R}$ but nowhere else
- ullet different propagators follow by evaluation of G_{ab} in different regions



Variational principle with effective dissipation

[Floerchinger, JHEP 1609, 099 (2016)]

• decompose inverse two-point function

$$P_{ab}(p) = P_{1,ab}(p) - is_{\mathsf{I}}(-u^{\mu}p_{\mu}) P_{2,ab}(p)$$

with $s_{\rm I}(\omega) = {\rm sign}({\rm Im}\ \omega)$

• in position space, replace

$$\begin{split} s_{\rm I} \left(-u^\mu p_\mu \right) &= {\rm sign} \left({\rm Im} (-u^\mu p_\mu) \right) \\ &\to {\rm sign} \left({\rm Im} \left(i u^\mu \frac{\partial}{\partial x^\mu} \right) \right) = {\rm sign} \left({\rm Re} \left(u^\mu \frac{\partial}{\partial x^\mu} \right) \right) = s_{\rm R} \left(u^\mu \frac{\partial}{\partial x^\mu} \right) \end{split}$$

 \bullet this symbol appears also in $\Gamma[\Phi]$

Retarded functional derivative

[Floerchinger, JHEP 1609, 099 (2016)]

 real and causal dissipative field equations follow from analytically continued effective action

$$\left. \frac{\delta \Gamma[\Phi]}{\delta \Phi_a(x)} \right|_{\rm ret} = \sqrt{g} J(x)$$

to calculate retarded variational derivative determine

$$\delta\Gamma[\Phi]$$

by varying the fields $\delta\Phi(x)$ including dissipative terms

set signs according to

$$s_{\mathsf{R}}(u^{\mu}\partial_{\mu}) \ \delta\Phi(x) \to -\delta\Phi(x), \qquad \qquad \delta\Phi(x) \ s_{\mathsf{R}}(u^{\mu}\partial_{\mu}) \to +\delta\Phi(x)$$

- proceed as usual
- opposite choice of sign: field equations for backward time evolution

Damped harmonic oscillator 1

equation of motion

$$m\ddot{x}+c\dot{x}+kx=0$$
 or with $\omega_0=\sqrt{k/m}$ and $\zeta=c/\sqrt{4mk}$
$$\ddot{x}+2\zeta\omega_0\dot{x}+\omega_0^2x=0$$

• is there an action for damped oscillator? This does not work:

$$\int \frac{d\omega}{2\pi} \, \frac{m}{2} x^*(\omega) \left[\omega^2 + 2i \, \omega \, \zeta \omega_0 - \omega_0^2 \right] x(\omega)$$

consider inverse propagator

$$\omega^2 + 2i\,s_{\rm I}(\omega)\,\omega\,\zeta\omega_0 - \omega_0^2$$

with sign function

$$s_{\mathsf{I}}(\omega) = \mathsf{sign}\left(\mathsf{Im}\,\omega\right)$$

Damped harmonic oscillator 2

effective action

$$\Gamma[x] = \int \frac{d\omega}{2\pi} \frac{m}{2} x^*(\omega) \left[-\omega^2 - 2i \, s_{\mathsf{I}}(\omega) \, \omega \, \zeta \omega_0 + \omega_0^2 \right] x(\omega)$$
$$= \int dt \left\{ -\frac{1}{2} m \dot{x}^2 + \frac{1}{2} c \, x \, s_{\mathsf{R}}(\partial_t) \dot{x} + \frac{1}{2} k x^2 \right\}$$

where the second line uses

$$s_{\mathsf{I}}(\omega) = \mathsf{sign}(\mathsf{Im}\,\omega) \to \mathsf{sign}(\mathsf{Im}\,i\partial_t) = \mathsf{sign}(\mathsf{Re}\,\partial_t) = s_{\mathsf{R}}(\partial_t)$$

variation gives up to boundary terms

$$\delta\Gamma = \int dt \left\{ m\ddot{x} \, \delta x + \frac{1}{2} c \, \delta x \, s_{\mathsf{R}}(\partial_t) \dot{x} - \frac{1}{2} c \, \dot{x} \, s_{\mathsf{R}}(\partial_t) \delta x + kx \, \delta x \right\}$$

- set now $s_{\mathsf{R}}(\partial_t)\delta x \to -\delta x$ and $\delta x \, s_{\mathsf{R}}(\partial_t) \to \delta x$. Defines $\frac{\delta \Gamma}{\delta x}|_{\mathsf{ret}}$.
- equation of motion for forward time evolution

$$\left. \frac{\delta \Gamma}{\delta x} \right|_{\text{ret}} = m\ddot{x} + c\dot{x} + kx = 0$$

$Entropy\ production$

[Floerchinger, JHEP 1609, 099 (2016)]

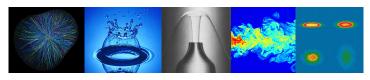
• analysis of general covariance leads to entropy production law

$$\nabla_{\mu}s^{\mu} = \frac{1}{\sqrt{g}}\frac{\delta\Gamma_{D}}{\delta\Phi_{a}}\Big|_{\rm ret}\beta^{\lambda}\partial_{\lambda}\Phi_{a} + \beta_{\mu}\nabla_{\nu}\left(-\frac{2}{\sqrt{g}}\frac{\delta\Gamma_{D}}{\delta g_{\mu\nu}}\Big|_{\rm ret}\right)$$

- should be positive by second law of thermodynamics
- so far only understood close-to-equilibrium
- . e.g. for viscous fluid

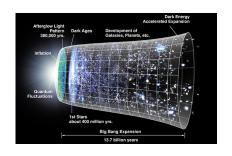
$$\nabla_{\mu} s^{\mu} = \frac{1}{T} \left[2\eta \sigma_{\mu\nu} \sigma^{\mu\nu} + \zeta (\nabla_{\rho} u^{\rho})^2 \right]$$

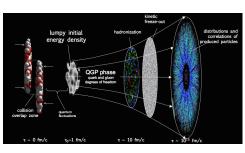
Fluid dynamics



- long distances, long times or strong enough interactions
- quantum fields form a fluid!
- needs macroscopic fluid properties
 - ullet equation of state $p(T,\mu)$
 - shear viscosity $\eta(T,\mu)$
 - bulk viscosity $\zeta(T, \mu)$
 - \bullet heat conductivity $\kappa(T,\mu)$
 - relaxation times, ...
- ab initio calculation of transport properties difficult but in principle fixed by microscopic properties encoded in lagrangian
- standard model of high energy nuclear collisions based on relativistic dissipative fluid dynamics
- ongoing experimental and theoretical effort to understand this better

$Big\ bang\ -\ little\ bang\ analogy$



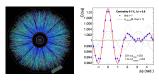


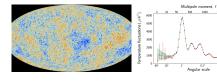
- cosmol. scale: MPc= 3.1×10^{22} m nuclear scale: fm= 10^{-15} m
- Gravity + QED + Dark sector
- one big event
 - t
- very many events
- dynamical description as a fluid
- all information must be reconstructed from final state

QCD

Fluid dynamic perturbation theory for heavy ions

[Floerchinger & Wiedemann, PLB 728, 407 (2014)] [ongoing work with E. Grossi, J. Lion, A. Mazeliauskas]





- goal: determine QCD fluid properties from experiments
- so far: numerical fluid simulations e.g. [Heinz & Snellings (2013)]
- new idea: solve fluid equations for smooth and symmetric background and order-by-order in perturbations
- less numerical effort more systematic studies
- good convergence properties [Floerchinger et al., PLB 735, 305 (2014), Brouzakis et al. PRD 91, 065007 (2015)]
- similar to cosmological perturbation theory

Dissipation in cosmology

[Floerchinger, Tetradis & Wiedemann, PRL 114, 091301 (2015)]

Evolution of energy density in first order viscous fluid dynamics

$$u^{\mu}\partial_{\mu}\epsilon + (\epsilon + p)\nabla_{\mu}u^{\mu} - \zeta\Theta^{2} - 2\eta\sigma^{\mu\nu}\sigma_{\mu\nu} = 0$$

with

- ullet bulk viscosity ζ
- ullet shear viscosity η

For $\vec{v}^2 \ll c^2$ and Newtonian potentials $\Phi, \Psi \ll 1$

$$\dot{\epsilon} + \vec{v} \cdot \vec{\nabla} \epsilon + (\epsilon + p) \left(3 \frac{\dot{a}}{a} + \vec{\nabla} \cdot \vec{v} \right)$$

$$= \frac{\zeta}{a} \left[3 \frac{\dot{a}}{a} + \vec{\nabla} \cdot \vec{v} \right]^2 + \frac{\eta}{a} \left[\partial_i v_j \partial_i v_j + \partial_i v_j \partial_j v_i - \frac{2}{3} (\vec{\nabla} \cdot \vec{v})^2 \right]$$

Fluid dynamic backreaction

[Floerchinger, Tetradis & Wiedemann, PRL 114, 091301 (2015)]

Expectation value of energy density $\bar{\epsilon} = \langle \epsilon \rangle$

$$\frac{1}{a}\dot{\bar{\epsilon}} + 3H\left(\bar{\epsilon} + \bar{p} - 3\bar{\zeta}H\right) = D$$

with dissipative backreaction term

$$D = \frac{1}{a^2} \langle \eta \left[\partial_i v_j \partial_i v_j + \partial_i v_j \partial_j v_i - \frac{2}{3} \partial_i v_i \partial_j v_j \right] \rangle$$
$$+ \frac{1}{a^2} \langle \zeta [\vec{\nabla} \cdot \vec{v}]^2 \rangle + \frac{1}{a} \langle \vec{v} \cdot \vec{\nabla} \left(p - 6\zeta H \right) \rangle$$

- D vanishes for unperturbed homogeneous and isotropic universe
- D has contribution from shear & bulk viscous dissipation and thermodynamic work done by contraction against pressure gradients
- dissipative terms in D are positive semi-definite
- for spatially constant viscosities and scalar perturbations only

$$D = \frac{\bar{\zeta} + \frac{4}{3}\bar{\eta}}{a^2} \int d^3q \ P_{\theta\theta}(q)$$

Dissipation of perturbations

[Floerchinger, Tetradis & Wiedemann, PRL 114, 091301 (2015)]

Dissipative backreaction does not need negative effective pressure

$$\frac{1}{a}\dot{\bar{\epsilon}} + 3H\left(\bar{\epsilon} + \bar{p}_{\text{eff}}\right) = D$$

- D is an integral over perturbations, could become large at late times.
- Can it potentially accelerate the universe?
- Need additional equation for scale parameter a
- Use trace of Einstein's equations $R=8\pi G_{\mathrm{N}}T^{\mu}_{\ \mu}$

$$\frac{1}{a}\dot{H} + 2H^2 = \frac{4\pi G_{\rm N}}{3} \left(\bar{\epsilon} - 3\bar{p}_{\rm eff}\right)$$

does not depend on unknown quantities like $\langle (\epsilon + p_{\rm eff}) u^\mu u^\nu \rangle$

• To close the equations one needs equation of state $\bar{p}_{\rm eff}=\bar{p}_{\rm eff}(\bar{\epsilon})$ and dissipation parameter D

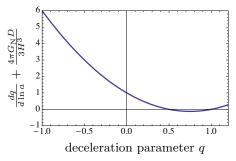
Deceleration parameter

[Floerchinger, Tetradis & Wiedemann, PRL 114, 091301 (2015)]

- ullet assume now vanishing effective pressure $ar{p}_{ ext{eff}}=0$
- \bullet obtain for deceleration parameter $q=-1-\frac{\dot{H}}{aH^2}$

$$-\frac{dq}{d \ln a} + 2(q-1)\left(q - \frac{1}{2}\right) = \frac{4\pi G_{\rm N} D}{3H^3}$$

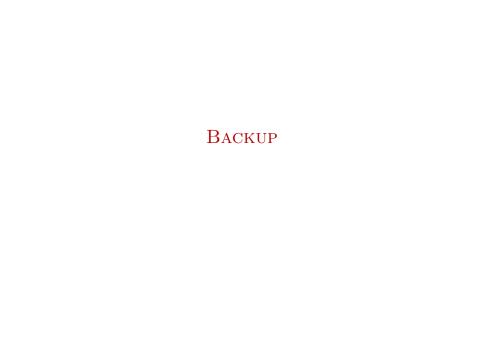
- for D=0 attractive fixed point at $q_*=\frac{1}{2}$ (deceleration)
- for D > 0 fixed point shifted towards $q_* < 0$ (acceleration)



Conclusions

- quantum field theory & information theory are entangled!
- could be essential element for universal non-equilibrium theory
- entanglement helps to understand "thermal effects" in e^+e^- and other collider experiments

 - at very early times theory effectively conformal $\frac{1}{\tau}\gg m,q$ entanglement entropy extensive in rapidity $\frac{dS}{d\Delta n}=\frac{c}{6}$
 - reduced density matrix for excitations at early times thermal $T=\frac{\hbar}{2\pi\tau}$
- experiments with cold atoms could allow to investigate entanglement directly
- effectively dissipative dynamics can have interesting consequences for cosmology



Coarse graining etc.

- entropy in quantum system can emerge when
 - system is divided into pieces with reduced density matrix
 - subsystems are composed again as mixed states
- cuts may divide
 - different regions
 - high-momentum and low-momentum
 - "system" and "bath"
- entropy in classical systems from coarse graining phase space
- entropy in kinetic theory from neglecting two-particle correlations (Boltzmann's "Stosszahlansatz")

Transverse coordinates

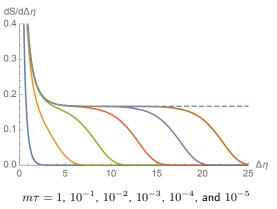
- So far dynamics strictly confined to 1+1 dimensions
- Transverse coordinates may fluctuate, can be described by Nambu-Goto action $(h_{\mu\nu}=\partial_{\mu}X^{m}\partial_{\nu}X_{m})$

$$\begin{split} S_{\text{NG}} &= \int d^2x \sqrt{-\text{det}\,h_{\mu\nu}}\,\left\{-\sigma + \ldots\right\} \\ &\approx \int d^2x \sqrt{g}\,\Big\{-\sigma - \frac{\sigma}{2}g^{\mu\nu}\partial_{\mu}X^i\partial_{\nu}X^i + \ldots\Big\} \end{split}$$

ullet Two additional, massless, bosonic degrees of freedom corresponding to transverse coordinates X^i with i=1,2.

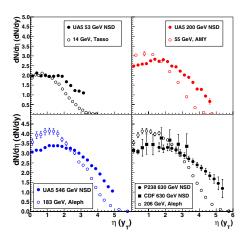
Free massive fermions

 \bullet Entanglement entropy can also be calculated for free Dirac fermions of mass m



- ullet Same universal plateau c/6 with c=1 at early time
- Conformal limit corresponds to non-interacting fermions
- Consistent with or without bosonization

Rapidity distribution



[open (filled) symbols: e⁺e⁻ (pp), Grosse-Oetringhaus & Reygers (2010)]

- ullet Rapidity distribution $dN/d\eta$ has plateau around midrapidity
- Only logarithmic dependence on collision energy

Experimental access to entanglement?

- Could longitudinal entanglement be tested experimentally?
- ullet Unfortunately entropy density $dS/d\eta$ not straight-forward to access.
- Measured in e^+e^- is the number of charged particles per unit rapidity $dN_{\rm ch}/d\eta$ (rapidity defined with respect to the thrust axis)
- Around mid-rapidity logarithmic dependence on the collision energy.
- \bullet Typical values for collision energies $\sqrt{s}=14-206~{\rm GeV}$ in the range

$$dN_{\rm ch}/d\eta \approx 2-4$$

 \bullet Entropy per particle S/N can be estimated for a hadron resonance gas in thermal equilibrium $S/N_{
m ch}=7.2$ would give

$$dS/d\eta \approx 14 - 28$$

 This is an upper bound: correlations beyond one-particle functions would lead to reduced entropy.

Temperature and entanglement entropy

- For conformal fields, entanglement entropy has also been calculated at non-zero temperature.
- For static interval of length l [Calabrese, Cardy (2004)]

$$S(T,l) = \frac{c}{3} \ln \left(\frac{1}{\pi T \epsilon} \sinh(\pi l T) \right) + \text{const}$$

Compare this to our result in expanding geometry

$$S(\tau,\Delta\eta) = \frac{c}{3} \ln \left(\frac{2\tau}{\epsilon} \sinh(\Delta\eta/2) \right) + \text{constant}$$

• Expressions agree for $l=\tau\Delta\eta$ (with metric $ds^2=-d\tau^2+\tau^2d\eta^2$) and time-dependent temperature

$$T = \frac{1}{2\pi\tau}$$

Alternative derivation: mode functions

• Fluctuation field $\varphi = \phi - \bar{\phi}$ has equation of motion

$$\partial_{\tau}^{2}\varphi(\tau,\eta) + \frac{1}{\tau}\partial_{\tau}\varphi(\tau,\eta) + \left(M^{2} - \frac{1}{\tau^{2}}\frac{\partial^{2}}{\partial\eta^{2}}\right)\varphi(\tau,\eta) = 0$$

• Solution in terms of plane waves

$$\varphi(\tau, \eta) = \int \frac{dk}{2\pi} \left\{ a(k)f(\tau, |k|)e^{ik\eta} + a^{\dagger}(k)f^{*}(\tau, |k|)e^{-ik\eta} \right\}$$

Mode functions as Hankel functions

$$f(\tau, k) = \frac{\sqrt{\pi}}{2} e^{\frac{k\pi}{2}} H_{ik}^{(2)}(M\tau)$$

or alternatively as Bessel functions

$$\bar{f}(\tau, k) = \frac{\sqrt{\pi}}{\sqrt{2\sinh(\pi k)}} J_{-ik}(M\tau)$$

$Bogoliubov\ transformation$

Mode functions are related

$$\begin{split} \bar{f}(\tau, k) = & \alpha(k) f(\tau, k) + \beta(k) f^*(\tau, k) \\ f(\tau, k) = & \alpha^*(k) \bar{f}(\tau, k) - \beta(k) \bar{f}^*(\tau, k) \end{split}$$

Creation and annihilation operators are related by

$$\bar{a}(k) = \alpha^*(k)a(k) - \beta^*(k)a^{\dagger}(k)$$
$$a(k) = \alpha(k)\bar{a}(k) + \beta(k)\bar{a}^{\dagger}(k)$$

Bogoliubov coefficients

$$\alpha(k) = \sqrt{\frac{e^{\pi k}}{2\sinh(\pi k)}} \qquad \beta(k) = \sqrt{\frac{e^{-\pi k}}{2\sinh(\pi k)}}$$

• Vacuum $|\Omega\rangle$ with respect to a(k) such that $a(k)|\Omega\rangle = 0$ contains excitations with respect to $\bar{a}(k)$ such that $\bar{a}(k)|\Omega\rangle \neq 0$ and vice versa

Role of different mode functions

- \bullet Hankel functions $f(\tau,k)$ are superpositions of positive frequency modes with respect to Minkowski time t
- ullet Bessel functions $ar{f}(au,k)$ are superpositions of positive and negative frequency modes with respect to Minkowski time t
- ullet At very early time $1/ au\gg M$ conformal symmetry

$$ds^2 = \tau^2 \left[-d\ln(\tau)^2 + d\eta^2 \right]$$

- ullet Hankel functions f(au,k) are superpositions of positive and negative frequency modes with respect to conformal time $\ln(au)$
- ullet Bessel functions $ar{f}(au,k)$ are superpositions of positive frequency modes with respect to conformal time $\ln(au)$

Occupation numbers

• Minkowski space coherent states have two-point functions

$$\langle \bar{a}^{\dagger}(k)\bar{a}(k')\rangle_{c} = \bar{n}(k) 2\pi \delta(k-k') = |\beta(k)|^{2} 2\pi \delta(k-k')$$
$$\langle \bar{a}(k)\bar{a}(k')\rangle_{c} = \bar{u}(k) 2\pi \delta(k+k') = -\alpha^{*}(k)\beta^{*}(k) 2\pi \delta(k+k')$$
$$\langle \bar{a}^{\dagger}(k)\bar{a}^{\dagger}(k')\rangle_{c} = \bar{u}^{*}(k) 2\pi \delta(k+k') = -\alpha(k)\beta(k) 2\pi \delta(k+k')$$

Occupation number

$$\bar{n}(k) = |\beta(k)|^2 = \frac{1}{e^{2\pi k} - 1}$$

 \bullet Bose-Einstein distribution with excitation energy $E=|k|/\tau$ and temperature

$$T = \frac{1}{2\pi\tau}$$

• Off-diagonal occupation number $\bar{u}(k) = -1/(2\sinh(\pi k))$ make sure we still have pure state

Local description

- \bullet Consider now rapidity interval $(-\Delta\eta/2,\Delta\eta/2)$
- Fourier expansion becomes discrete

$$\varphi(\eta) = \frac{1}{L} \sum_{n = -\infty}^{\infty} \varphi_n \ e^{in\pi \frac{\eta}{\Delta \eta}}$$

$$\varphi_n = \int_{-\Delta\eta/2}^{\Delta\eta/2} d\eta \ \varphi(\eta) \ \frac{1}{2} \left[e^{-in\pi \frac{\eta}{\Delta\eta}} + (-1)^n e^{in\pi \frac{\eta}{\Delta\eta}} \right]$$

Relation to continuous momentum modes by integration kernel

$$\varphi_n = \int \frac{dk}{2\pi} \sin(\frac{k\Delta\eta}{2} - \frac{n\pi}{2}) \left[\frac{1}{k - \frac{n\pi}{\Delta\eta}} + \frac{1}{k + \frac{n\pi}{\Delta\eta}} \right] \varphi(k)$$

Local density matrix determined by correlation functions

$$\langle \varphi_n \rangle$$
, $\langle \pi_n \rangle$, $\langle \varphi_n \varphi_m \rangle_c$, etc.

Emergence of locally thermal state

• Mode functions at early time

$$\bar{f}(\tau, k) = \frac{1}{\sqrt{2k}} e^{-ik\ln(\tau) - i\theta(k, M)}$$

• Phase varies strongly with k for $M \to 0$

$$\theta(k, M) = k \ln(M/2) + \arg(\Gamma(1 - ik))$$

ullet Off-diagonal term $ar{u}(k)$ have factors strongly oscillating with k

$$\langle \varphi(\tau, k) \varphi^*(\tau, k') \rangle_c = 2\pi \delta(k - k') \frac{1}{|k|} \times \left\{ \left[\frac{1}{2} + \bar{n}(k) \right] + \cos\left[2k \ln(\tau) + 2\theta(k, M)\right] \bar{u}(k) \right\}$$

cancel out when going to finite interval!

 \bullet Only Bose-Einstein occupation numbers $\bar{n}(k)$ remain

Entanglement and deep inelastic scattering

- How strongly entangled is the nuclear wave function?
- What is the entropy of quasi-free partons and can it be understood as a result of entanglement? [Kharzeev, Levin (2017)]

$$S = \ln[xG(x)]$$

- Does saturation at small Bjorken-x have an entropic meaning?
- Entanglement entropy and entropy production in the color glass condensate [Kovner, Lublinsky (2015)]
- Could entanglement entropy help for a non-perturbative extension of the parton model?
- Entropy of perturbative and non-perturbative Pomeron descriptions [Shuryak, Zahed (2017)]