

Mode-by-mode fluid dynamics for relativistic heavy ion collisions

Stefan Flörchinger (CERN)

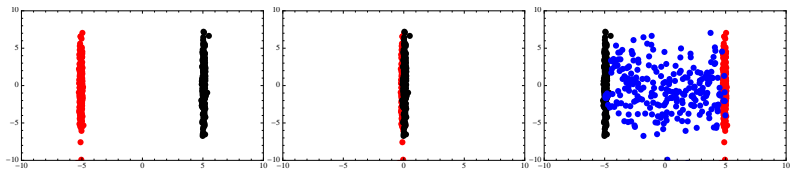
Ohio State University, April 10, 2014

based on work with Urs A. Wiedemann, Andrea Beraudo et al.

- Fluctuations around Bjorken Flow and the onset of turbulent phenomena, [[JHEP 11, 100 \(2011\)](#), with U. A. Wiedemann]
- Mode-by-mode fluid dynamics for relativistic heavy ion collisions [[Phys. Lett. B, 728, 407 \(2014\)](#), with U. A. Wiedemann]
- Characterization of initial fluctuations for the hydrodynamical description of heavy ion collisions, [[Phys. Rev. C 88, 044906 \(2013\)](#), with U. A. Wiedemann]
- Kinetic freeze-out, particle spectra and harmonic flow coefficients from mode-by-mode hydrodynamics, [[arXiv:1311.7613](#), with U. A. Wiedemann]
- How (non-) linear is the hydrodynamics of heavy ion collisions? [[arXiv:1312.5482](#), with U. A. Wiedemann, A. Beraudo, L. Del Zanna, G. Inghirami, V. Rolando]
- Mode-by-mode hydrodynamics: ideas and concepts [[arXiv:1401.2339](#)]

Introduction

Heavy Ion Collisions



- ions are strongly Lorentz-contracted
- *some* medium is produced after collision
- medium expands in longitudinal direction and gets diluted

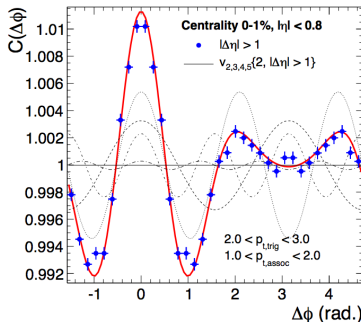
Evolution in time

- Non-equilibrium evolution at early times
 - initial state at from QCD? Color Glass Condensate? ...
 - thermalization via strong interactions, plasma instabilities, particle production, ...
- Local thermal and chemical equilibrium
 - strong interactions lead to short thermalization times
 - evolution from relativistic fluid dynamics
 - expansion, dilution, cool-down
- Chemical freeze-out
 - for small temperatures one has mesons and baryons
 - inelastic collision rates become small
 - particle species do not change any more
- Thermal freeze-out
 - elastic collision rates become small
 - particles stop interacting
 - particle momenta do not change any more

Fluid dynamic regime

- Assumes strong interaction effects leading to local equilibrium.
- Fluid dynamic variables
 - thermodynamic variables: e.g. $T(x)$, $\mu(x)$,
 - fluid velocity $u^\mu(x)$,
 - shear stress tensor $\pi^{\mu\nu}(x)$,
 - bulk viscous pressure $\pi_{\text{Bulk}}(x)$.
- Can be formulated as derivative expansion for $T^{\mu\nu}$.
- Hydrodynamics is universal: many details of microscopic theory not important.
- Some macroscopic properties are important:
 - ideal hydro: needs equation of state $p = p(T, \mu)$, $n = n(T, \mu)$ from thermodynamics
 - first order hydro: needs also transport coefficients like shear viscosity $\eta = \eta(T, \mu)$ and bulk viscosity $\zeta(T, \mu)$ from linear response theory
 - second order hydro: needs also relaxation times τ_{Shear} , τ_{Bulk} etc.

Experimental proof for fluctuations: v_3 and v_5



(ALICE 2011, similar pictures also from CMS, ATLAS, Phenix)

- One can expand two-point correlation function

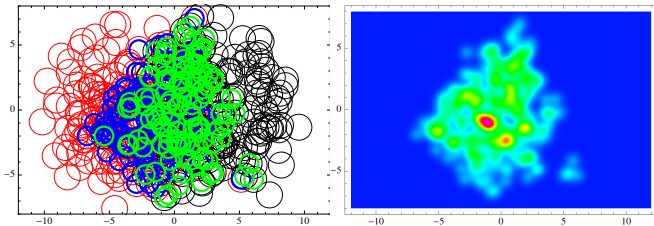
$$C(\Delta\phi) \sim 1 + \sum_{n=2}^{\infty} 2 v_n \cos(n \Delta\phi)$$

- Without fluctuations one would expect from mirror symmetry

$$v_3 = v_5 = \dots = 0.$$

Event-by-event fluctuations

- Argument for $v_3 = v_5 = 0$ is based on smooth and symmetric energy density distribution.
- Deviations from this can come from event-by-event fluctuations.
- One example is Glauber model

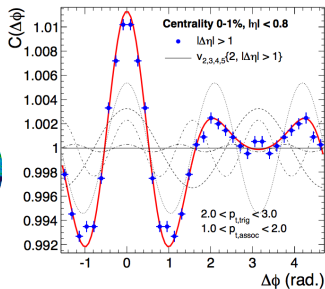
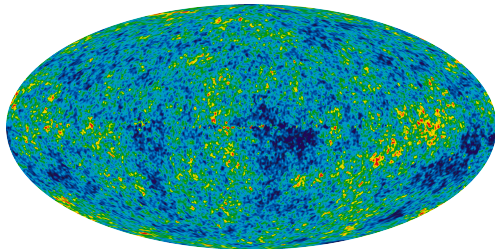


- The initial transverse density distribution fluctuates event-by-event and this leads to sizeable v_3 and v_5 .
- More generally also other initial hydro fields may fluctuate: fluid velocity, shear stress, baryon number density etc.

What fluctuations are interesting and why?

- **Initial hydro fluctuations:** Event-by-event perturbations around the average of hydrodynamical fields at time τ_0 :
 - energy density ϵ
 - fluid velocity u^μ
 - shear stress $\pi^{\mu\nu}$
 - more general also: baryon number density n_B , electric charge density, electromagnetic fields, ...
- measure for deviations from equilibrium
- contain interesting information from early times
- governed by universal evolution equations
- can be used to constrain thermodynamic and transport properties

Similarities to cosmic microwave background



- fluctuation spectrum contains info from early times
- many numbers can be measured and compared to theory
- can lead to detailed understanding of evolution and properties
- could trigger precision era in heavy ion physics

A program to understand fluctuations

- ➊ Initial fluctuations at initialization time of hydro should be characterized and quantified completely.
- ➋ Fluctuations have to be propagated through the hydrodynamical regime.
- ➌ Contribution of different fluctuations to the particle spectra must be understood and quantified.
- ➍ Fluctuations generated from non-hydro sources (such as jets) have to be taken into account.

Background-fluctuation splitting

- Background or average over many events is described by smooth fields

$$w_0 = \langle w \rangle$$
$$u_0^\mu = \langle u^\mu \rangle$$

- Fluctuations are added on top

$$w = w_0 + w_1$$
$$u^\mu = u_0^\mu + u_1^\mu$$

- For background one may assume Bjorken boost and azimuthal rotation invariance

$$w_0 = w_0(\tau, r)$$
$$u_0^\mu = (u_0^\tau(\tau, r), u_0^r(\tau, r), 0, 0)$$

Characterization of initial conditions

Characterization of transverse density 1

Fluctuations in initial transverse enthalpy density $w(r, \phi)$ can be characterized in terms of eccentricities $\epsilon_{n,m}$ and angles $\psi_{n,m}$
[Ollitrault, Teaney, Yan, Luzum, and others]

$$\epsilon_{n,m} e^{im\psi_{n,m}} = \frac{\int dr \int_0^{2\pi} d\varphi r^{n+1} e^{im\varphi} w(r, \varphi)}{\int dr \int_0^{2\pi} d\varphi r^{n+1} w(r, \varphi)}$$

- $w(r, \phi)$ completely determined by set of all $\epsilon_{n,m}$ and $\psi_{n,m}$
- closely related method is based on cumulants [Teaney, Yan]
- no positive transverse density can be associated to small set of cumulants (beyond Gaussian order) such that higher order cumulants vanish
- generalization to velocity and shear fluctuations not known

Characterization of transverse density 2

Characterization based on Bessel-Fourier expansion [Coleman-Smith, Petersen & Wolpert, 2012]

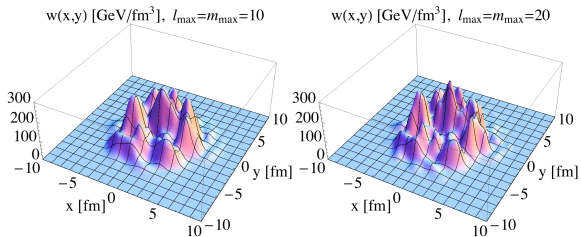
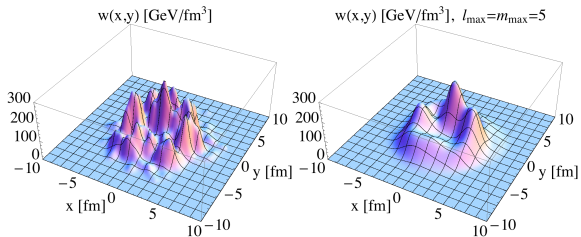
$$w(r, \phi) = \sum_{m,n} A_{m,n} e^{im\phi} J_m(\lambda_{m,n} \frac{r}{r_0})$$

Characterization based on Bessel-Fourier expansion and background density [Floerchinger & Wiedemann, 2013]

$$w(r, \phi) = w_{\text{BG}}(r) + w_{\text{BG}}(r) \sum_{m=-m_{\text{max}}}^{m_{\text{max}}} \sum_{l=1}^{l_{\text{max}}} \tilde{w}_l^{(m)} e^{im\phi} J_m(k_l^{(m)} r)$$

- $w(r, \phi)$ completely determined by set of all $\tilde{w}_l^{(m)}$
- higher l correspond to smaller spatial resolution
- single or few coefficients $\tilde{w}_l^{(m)}$ lead to *positive density*
- single modes can be propagated in hydro
- works similar for vectors (velocity) and tensors (shear stress)

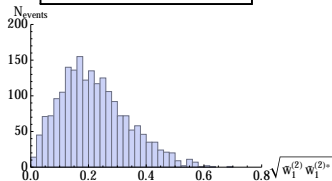
Transverse density from Glauber model



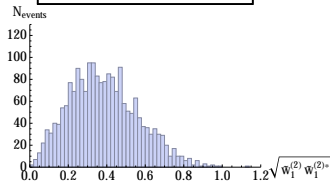
Distribution of weights

From Monte-Carlo Glauber model. Some examples:

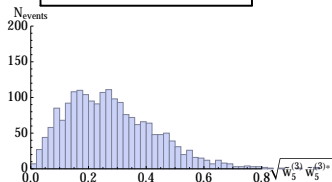
2000 events, $b=2\text{fm}$, $m=2$, $l=1$



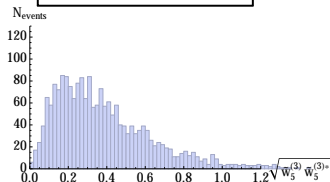
2000 events, $b=4\text{ fm}$, $m=2$, $l=1$



2000 events, $b=2\text{fm}$, $m=3$, $l=5$



2000 events, $b=4\text{ fm}$, $m=3$, $l=5$



Velocity fluctuation

- Initial velocity fluctuations at $\tau_0 \approx 0.5 \text{ fm}/c$ are conceivable.
- Characterization similar as for density fluctuations. Two polarizations

$$u^r = u_{\text{BG}}^r + \frac{1}{\sqrt{2}}(\tilde{u}^- + \tilde{u}^+)$$

$$u^\phi = \frac{i}{\sqrt{2}r}(\tilde{u}^- - \tilde{u}^+),$$

with

$$\tilde{u}^-(r, \phi) = \sum_{m,l} \tilde{u}_l^{-(m)} e^{im\phi} J_{m-1} \left(k_l^{(m)} r \right)$$

$$\tilde{u}^+(r, \phi) = \sum_{m,l} \tilde{u}_l^{+(m)} e^{im\phi} J_{m+1} \left(k_l^{(m)} r \right).$$

- Would be interesting to search for them in experimental data.

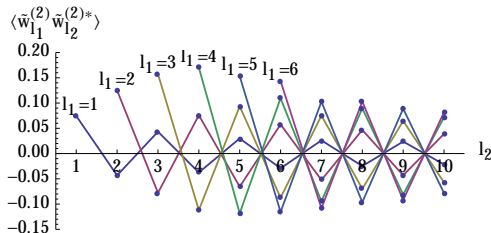
Event ensembles

- Event ensembles can be characterized in terms of functional probability distribution $p_{\tau_0}[w, u^\mu, \pi^{\mu\nu}, \dots]$.
- Simplest case is Gaussian form

$$p_{\tau_0} \sim \exp \left[-\frac{1}{2} \sum_{m=-m_{\max}}^{m_{\max}} \sum_{l_1, l_2=1}^{l_{\max}} T_{l_1 l_2}^{(m)} \tilde{w}_{l_1}^{(m)*} \tilde{w}_{l_2}^{(m)} \right]$$

- Fully determined by correlator

$$(T^{(m)})_{l_1 l_2}^{-1} = \langle \tilde{w}_{l_1}^{(m)} \tilde{w}_{l_2}^{(m)*} \rangle$$



Hydrodynamic evolution

Perturbative expansion

Write the hydrodynamic fields $h = (w, u^\mu, \pi^{\mu\nu}, \pi_{\text{Bulk}}, \dots)$

- at initial time τ_0 as

$$h = h_0 + \epsilon h_1$$

with h_0 the Background and ϵh_1 the fluctuation part

- at some later time $\tau > \tau_0$ as

$$h = h_0 + \epsilon h_1 + \epsilon^2 h_2 + \epsilon^3 h_3 + \dots$$

Solve for time evolution in this scheme

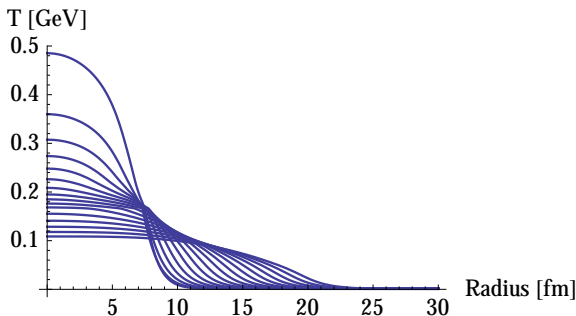
- h_0 is a solution of the full non-linear hydro equations but with higher symmetry: azimuthal rotation and Bjorken boost invariance
- h_1 is a solution of the linearized hydro equations around h_0 , can be solved mode-by-mode
- h_2 can be obtained by from interactions between modes etc.

Background evolution

System of coupled 1 + 1 dimensional non-linear partial differential equations for

- enthalpy density $w(\tau, r)$ or temperature $T(\tau, r)$
- fluid velocity $u^\tau(\tau, r), u^r(\tau, r)$
- two independent components of shear stress $\pi^{\mu\nu}(\tau, r)$

Can be easily solved numerically

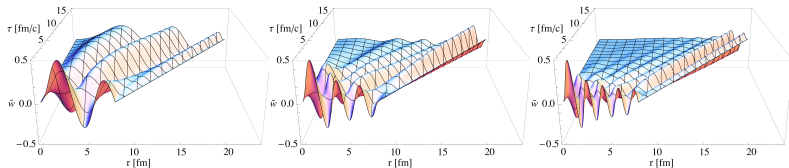


Evolving fluctuation modes

Linearized hydro equations are system of coupled 3 + 1 dimensional, linear partial differential equations. Use Fourier expansion

$$h_j(\tau, r, \phi, \eta) = \sum_m \int \frac{dk_\eta}{2\pi} h_j^{(m)}(\tau, r, k_\eta) e^{i(m\phi + k_\eta \eta)}.$$

This gives 1 + 1 dimensional linear partial differential equations that can be solved again numerically for initial conditions corresponding to each Bessel-Fourier mode.



Mode interactions

- Non-linear terms in the evolution equations for fluctuating fields lead to mode interaction terms of quadratic and higher order in the initial fluctuation fields.
- One can determine these terms from an iterative solution but that has not been fully worked out yet.
- The whole picture can be tested with complete numerical solution of the full hydro equations.

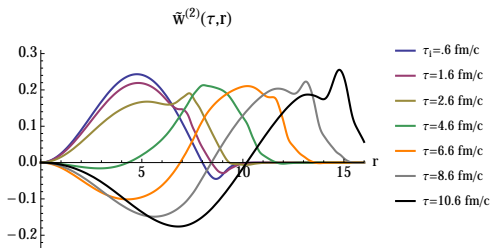
Scaling tests

- Start with single enthalpy density mode ($m = 2, l = 1$) on top of background

$$w(\tau_0, r, \phi) = w_{\text{BG}}(\tau_0, r) \left[1 + 2 \tilde{w}_1^{(2)} J_2(k_1^{(2)} r) \cos(2\phi) \right].$$

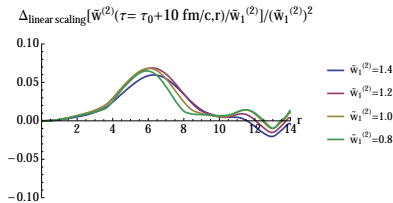
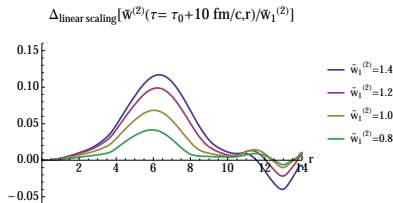
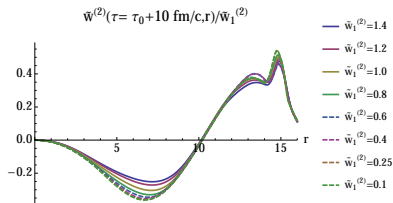
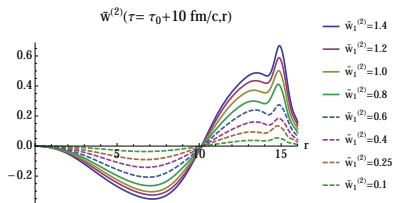
- Evolve this with hydro solver ECHO-QGP [Del Zanna *et al.*, EPJC 73, 2524 (2013)] and determine Fourier component

$$\tilde{w}^{(2)}(\tau, r) = w^{(2)}(\tau, r) / w_{\text{BG}}(\tau, r).$$



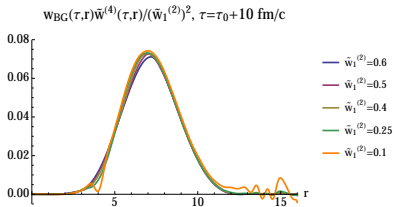
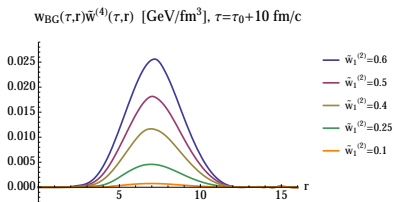
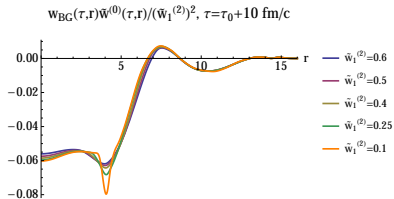
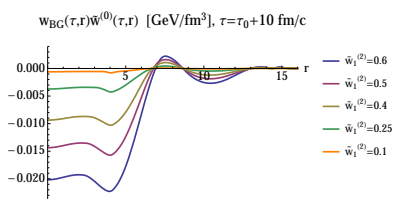
Scaling tests at first order

Compare enthalpy $\tilde{w}^{(2)}(\tau, r)$ at fixed τ for different initial weights



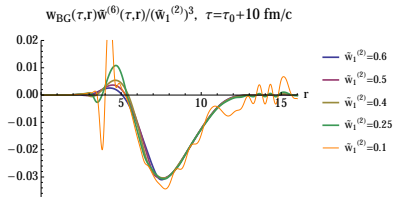
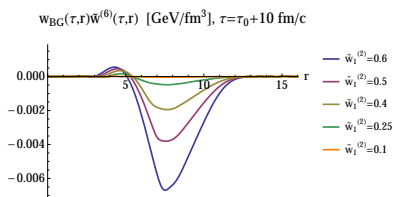
Scaling tests at second order

From symmetry considerations one expects that modes with $m = 0$ and $m = 4$ receive mainly quadratic contributions $\sim (\tilde{w}_1^{(2)})^2$.



Scaling tests at third order

From symmetry considerations one expects that modes $m = 6$ receive mainly cubic contributions $\sim (\tilde{w}_1^{(2)})^3$.

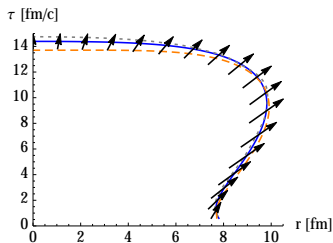


- Hydrodynamic response to initial enthalpy density perturbations perturbative.
- Non-linearities can be understood order-by-order and lead to characteristic “overtones”.
- Results motivate a more thorough development of hydrodynamic perturbation theory.

Kinetic freeze-out

Freeze-out surface

- Background and fluctuations are propagated until $T_{fo} = 120$ MeV is reached.
- Distribution functions are determined and free streaming is assumed for later times [Cooper & Frye].
- Background-fluctuation splitting and expansion in powers of perturbations can be used also at freeze-out.
- Freeze-out surface is azimuthally symmetric as background [Floerchinger, Wiedemann 2013].
- Generalization to kinetic hadronic scattering and decay phase possible.



(solid: $\eta/s = 0.08$, dotted: $\eta/s = 0$, dashed: $\eta/s = 0.3$)

Contribution of modes to “single event spectrum”

Particle spectrum (or its logarithm) can be expanded in contribution from different modes. To linear order:

$$\ln \left(\frac{dN^{\text{single event}}}{p_T dp_T d\phi dy} \right) = \underbrace{\ln S_0(p_T)}_{\text{from background}} + \underbrace{\sum_{m,l} \tilde{w}_l^{(m)} e^{im\phi} \theta_l^{(m)}(p_T)}_{\text{from fluctuations}}$$

- Each mode has it's own angle $\tilde{w}_l^{(m)} = |\tilde{w}_l^{(m)}| e^{im\psi_l^{(m)}}$.
- p_T -dependence of different modes described by $\theta_l^{(m)}(p_T)$.
- To quadratic order this gets supplemented by

$$\sum_{m_1, m_2, l_1, l_2} \tilde{w}_{l_1}^{(m_1)} \tilde{w}_{l_2}^{(m_2)} e^{i(m_1+m_2)\phi} \kappa_{l_1, l_2}^{(m_1, m_2)}(p_T).$$

- The non-linearities encoded in $\kappa_{l_1, l_2}^{(m_1, m_2)}(p_T)$ arise both from hydro evolution and from kinetic freeze-out itself.

Harmonic flow coefficients

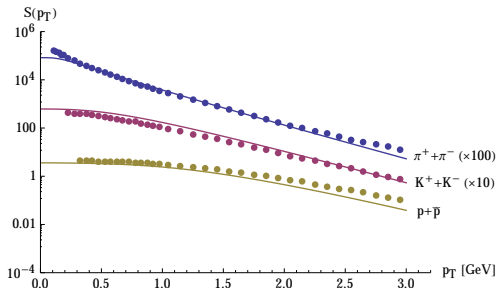
Double differential harmonic flow coefficient to lowest order

$$v_m^2\{2\}(p_T^a, p_T^b) = \sum_{l_1, l_2=1}^{l_{\max}} \theta_{l_1}^{(m)}(p_T^a) \theta_{l_2}^{(m)}(p_T^b) \langle \tilde{w}_{l_1}^{(m)} \tilde{w}_{l_2}^{(m)*} \rangle$$

- intuitive matrix expression
- in general no factorization
- higher order corrections important for non-central collisions

One-particle spectrum

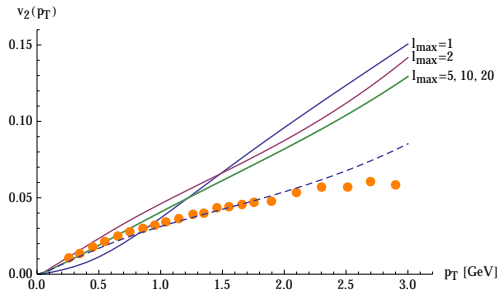
$$S(p_T) = dN/(2\pi p_T dp_T d\eta d\phi)$$



Points: 5% most central collisions, ALICE [[PRL 109, 252301 \(2012\)](#)]
Curves: Our calculation, no hadron rescattering and decays after freeze-out.

Harmonic flow coefficients for central collisions

Elliptic flow for charged particles



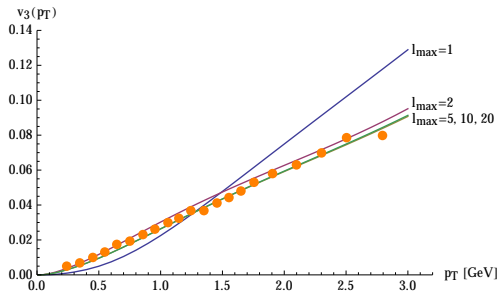
Points: 2% most central collisions, ALICE [\[PRL 107, 032301 \(2011\)\]](#)

Solid curves: Different maximal resolution l_{\max}

Dashed curve: Mode $(m=2, l=1)$ suppressed by factor 0.7

Harmonic flow coefficients for central collisions

Triangular flow for charged particles

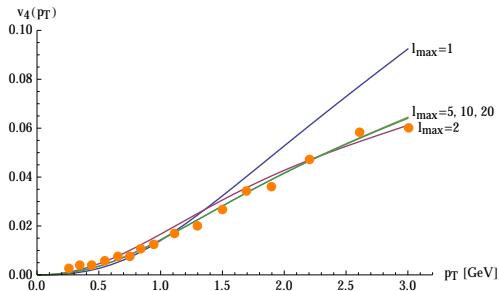


Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)]

Curves: Different maximal resolution l_{\max}

Harmonic flow coefficients for central collisions

Flow coefficient v_4 for charged particles

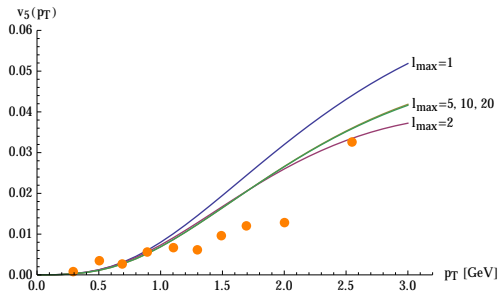


Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)]

Curves: Different maximal resolution l_{\max}

Harmonic flow coefficients for central collisions

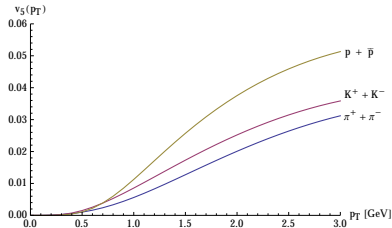
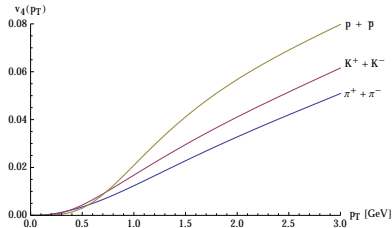
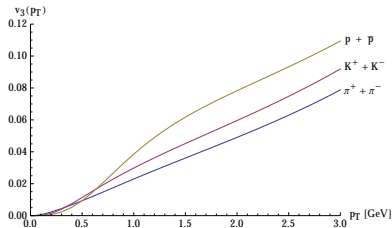
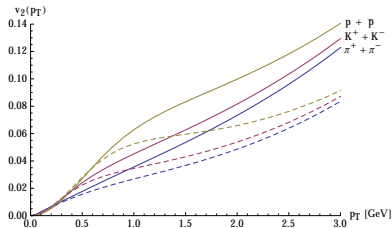
Flow coefficient v_5 for charged particles



Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)]

Curves: Different maximal resolution l_{\max}

Harmonic flow coefficients, central, particle identified



Summary and Conclusions

Conclusions

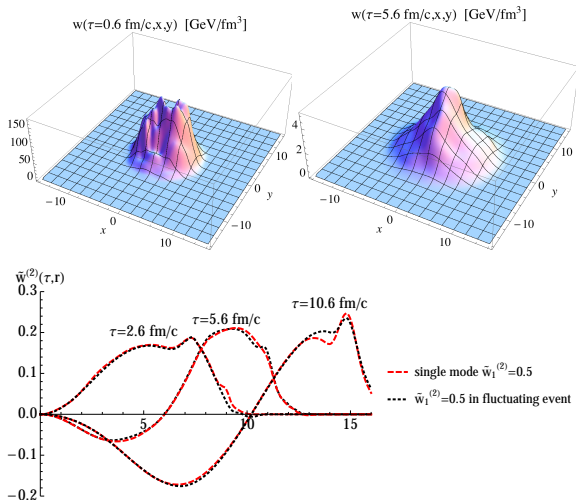
- Mode-by-mode perturbative hydrodynamic picture allows to determine response to initial density perturbations.
- Hydrodynamic evolution can be disentangled from initial state model.
- First study for enthalpy density fluctuations in Glauber model
 - yields good description of $v_m(p_T)$ for central collisions,
 - shows that fluctuations up to $l_{\max} \approx 5$ can be resolved.
- Fluctuations to be studied:

| | transverse plane | rapidity direction |
|----------------------------|------------------|--------------------|
| enthalpy density / entropy | ✓ | - |
| fluid velocity | - | - |
| shear stress | - | - |
| baryon number density | - | - |
| electromagnetic fields | - | - |
| electric charge density | - | - |
| chiral order parameter | - | - |

Backup

Scaling tests embedded in realistic event

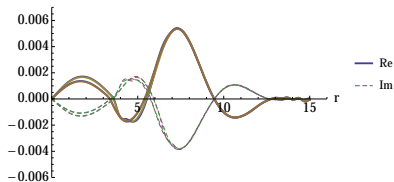
Embed mode ($m = 2, l = 1$) into realistic fluctuating event and compare to embedding into pure background.



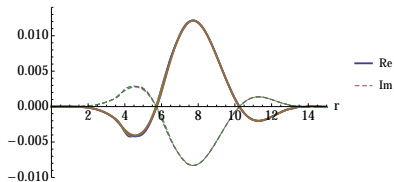
Scaling tests with several initial modes

Start with linear combination of $(m = 2, l = 2)$ and $(m = 3, l = 1)$ modes and test scaling for $m = 1$ and $m = 5$ response.

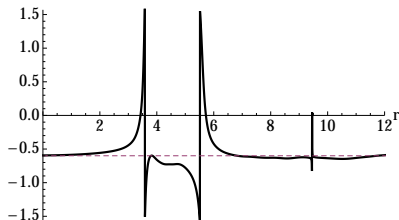
$$w_{BG}(\tau, r) \tilde{w}^{(1)}(\tau, r) / |\tilde{w}_2^{(2)} \tilde{w}_1^{(3)}| \text{ [GeV/fm}^3\text{]}$$



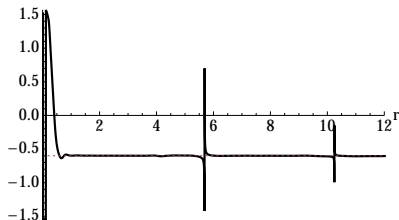
$$w_{BG}(\tau, r) \tilde{w}^{(5)}(\tau, r) / |\tilde{w}_2^{(2)} \tilde{w}_1^{(3)}| \text{ [GeV/fm}^3\text{]}$$



$$\text{Arg}[\tilde{w}^{(1)}(\tau, r)] \text{ for } \psi^{(2)}=0.0, \psi^{(3)}=-0.2$$



$$\text{Arg}[\tilde{w}^{(5)}(\tau, r)] \text{ for } \psi^{(2)}=0.0, \psi^{(3)}=-0.2$$



Generalized Glauber model

- Fluctuations due to nucleon positions: used so far

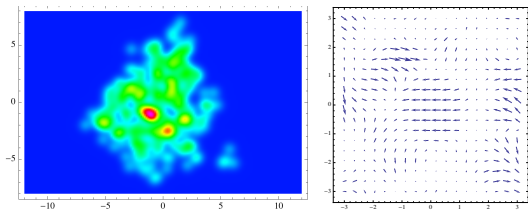
$$\epsilon(\tau, \mathbf{x}, y) = \sum_{i=1}^{N_{\text{part}}} \epsilon_w(\tau, \mathbf{x} - \mathbf{x}_i, y), \quad u^\mu = (1, 0, 0, 0)$$

- can be generalized to include also velocity fluctuations

$$T^{\mu\nu}(\tau, \mathbf{x}, y) = \sum_{i=1}^{N_{\text{part}}} T_w^{\mu\nu}(\tau, \mathbf{x} - \mathbf{x}_i, y)$$

- More generally describe primordial fluid fields by
 - expectation values $\langle \epsilon(\tau_0, \mathbf{x}, y) \rangle, \langle u^\mu(\tau_0, \mathbf{x}, y) \rangle, \langle n_B(\tau_0, \mathbf{x}, y) \rangle$
 - correlation functions $\langle \epsilon(\tau_0, \mathbf{x}, y) \epsilon(\tau_0, \mathbf{x}', y') \rangle$, etc.
- Origin of this fluctuations is initial state physics and early-time, non-equilibrium dynamics.

Velocity fluctuations



- also the velocity field will fluctuate at the initialization time τ_0
- take here transverse velocity for every participant to be Gaussian distributed with width $0.1c$
- vorticity $|\partial_1 u^2 - \partial_2 u^1|$ and divergence $|\partial_1 u^1 + \partial_2 u^2|$

